Practical Clustered Shading

Part 2/4

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What’s new?

- Preaching the Clustered gospel for two years
  - Nordic Game 2013
  - SIGGRAPH 2013
  - CEDEC 2013
  - SIGGRAPH Asia 2014

- Avalanche Studios still using it in production
  - Used in newly announced Just Cause 3
  - Very happy with it
  - All old lighting paths removed

- Interest in Clustered Shading increasing
  - Intel samples for PC and Android [Intel 14]
  - Gets name-dropped a lot by game developers
    - Although few have actually implemented it yet
Agenda

- History of lighting in the Avalanche Engine
- Why Clustered Shading?
- Adaptations for the Avalanche Engine
- Performance
- Future work
Just Cause 1 had 3 global pointlights. This meant that if, for instance, three streetlights were enabled and you fired your gun, one of the lights would shut off for the duration of the gun flash. Clearly, this solution was hardly ideal.

For Just Cause 2 we switched to a world-space 2D tiled solution where light indexes were stored in texels. The technique has been described in detail in the article ”Making it Large, Beautiful, Fast, and Consistent: Lessons Learned Developing Just Cause 2” in GPU Pro. This technique was actually in some ways similar to clustered shading, although much more limited and designed around DX9 level hardware. It worked reasonably well on platforms with decent dynamic branching, such as PC and Xenon, whereas the PS3 struggled. Ultimately this caused us to implement numerous workarounds to get PS3 running well, so that in the end this technique mostly ended up being a fallback option if the light count was too high for a specialized shader to work. The amount of specialized shaders also became quite a bit of a maintenance problems, and figuring out the light count a performance issue on the CPU side.
After Just Cause 2 we ended going the deferred shading route, initially using classic deferred. This worked relatively well for last generation console hardware and allowed us to support many more lights, different light types, shadow casting dynamic lights etc. This was great, but naturally we also got all the downsides of deferred shading, such as problems with transparency, problems with custom material or lighting models, as well as large increase in memory consumption. Initially we supported MSAA, but ultimately we dropped it in favor of FXAA for performance and memory reasons.

Unfortunately, the old forward pass also had to stick around for transparency to work to some extent, although it only ever supported pointlights and the lighting didn’t quite match the much more sophisticated deferred pass. For Mad Max we ultimately moved away from supporting transparency with lighting because of its problems with deferred, plus that the game environment has very little need for transparency anyway beyond particle effects. But for other projects where transparency might be desirable we started looking into alternatives, especially with a new generation consoles on the horizon at the time.
For Just Cause 3, which is next-gen/PC only, we went with Clustered Shading as our main lighting solution. For this kind of game it really wasn’t a feasible solution to drop transparency. We are still using deferred, but with clustered shading we can use the same lighting data for doing lighting in the forward passes for transparent objects. A nice bonus of having properly working transparency is that we could now use the Wire AA technique we invented. We have attempted to add this to Mad Max as well, but it has been much more problematic in a classic deferred context, so it’s questionable at this point if we will use it to any greater extent in that game. For Just Cause 3 it mostly just worked out of the box.
Tiled Deferred Shading and Forward+ (Tiled Forward Shading) are production proven and has shipped in real games, but they come with a bunch of drawbacks. Tiled deferred offers better performance than classic deferred, but doesn’t really solve any of our problems since all drawbacks of classic deferred stays around. In addition, it also imposes a new restriction in that all lights, and consequently shadow buffers, are now required up-front. However, this is a property it shares with all other techniques, including Forward+ and Clustered Shading, and even our old forward solution from JC2.

Forward+ has the advantage of working well with MSAA without hassles; however, while it can be made to work with transparency, it requires an extra pass, including another round of pre-z. The requirement of a full pre-z pass for this technique to work made this a non-starter for us. We didn’t bother implementing it for evaluation purposes as a full pre-z pass is not an option for us. We did at one point have a fairly complete pre-z pass in Just Cause 2, but over the development the pre-z pass was continuously trimmed until very little remained. The additional overhead just didn’t pay off, and the large increase in draw-call count was problematic. After we got a decent occlusion culling system in place there were very few cases pre-z did not, in fact, result in a performance drop. Pre-z is now only enabled on a handful of things specifically marked for pre-z by content creators, and a few code-driven systems that need it for other reasons.

Clustered Shading has the advantage of not requiring a pre-z pass, even in its forward incarnation, while working well with MSAA and transparency out of the box with no particular tricks or hacks. It has at the point of this writing to our knowledge not shipped in any real games so far, but it has been in production at Avalanche Studios since January 2013 and has so far worked really well for us and we expect it to make it all the way to shipping.
Clustered Shading is really decoupled from the choice between deferred or forward rendering. It works with both, so you’re not locked into one or the other. This way you can make an informed choice between the two approaches based on other factors, such as whether you need custom materials and lighting models, or need deferred effects such as screen-space decals, or simply based on performance.

The two tiled solutions need quite a bit of massaging to work reasonable well in all situations, especially with large amounts of depth discontinuities. There are proposed solutions that mitigate the problem, such as 2.5D culling, but they further complicate the code. For Clustered Shading it just falls out automatically and depth discontinuities do not cause performance problems. This allows Clustered Shading to maintain a more stable frame-rate regardless of scene depth complexity.
I will illustrate the point using a random screenshot from Just Cause 3. Now this isn’t a hand picked screenshot to show off the worst case, in fact, I wasn’t able to pick a screenshot myself. This was hand-picked by marketing for being awesome. But even so, it’s representative of what you can expect in the Just Cause series and really shows that this is a real problem in real games, and certainly so in the games that we make.
Here a number of large depth differences have been manually painted over the image to illustrate where you might expect a problem for tiled shading techniques. As you can see, they are fairly common and affect a fairly large part of the screen. One source of pain that’s not too well illustrated here, is vegetation, which tends to create lots of nasty depth discontinuities. There are some forests here, but for ground level gameplay you can certainly expect much more of that problem.
Now, if you thought the previous image looked bad, now put that into the context of an actual tiled setup. Here I have illustrated all the tiles that would be affected by the problem, and as you can see, a quite large percentage of the screen suffers from suboptimal lighting from depth discontinuities.
Let’s discuss the problem of depth discontinuities and illustrate how clustered shading solves it. Here’s a sample frustum with some depth values, including a few discontinuities.
Clustering and depth

- Tiled frustum

Here we added the tiles.
And this is the depth ranges you would get for a plain tiled shading algorithm. Clearly some ranges are fairly large.
With 2.5D culling the situation is notably improved. Now lights in the discontinuity area is not included. However, we do pay the full cost lights at both ends for both sides of the discontinuity. Also note that one very long depth range remains. This is because it’s not discontinuous, it’s a continuous slope. This situation would happen if you look down a hallway, or the ground plane, or moderate large surface at a grazing angle.
Now let’s look at a clustered frustum.
These are the depth ranges that we will need to consider. Note that we are paying for exactly one cluster’s depth at any given point.
If we go to explicit cluster bounds, the situation is even further improved, although in practice there may not be a huge difference between a fairly small range and an even smaller range, depending on the typical size of light sources.
Here we see the improvement from implicit bounds to explicit.
And here all techniques are compared. As you can see, explicit clustered is always the tightest. However, there are definitively areas where tiled with 2.5D culling is tighter than implicit cluster bounds. So in scenes with little depth complexity tiled could very well be faster. However, implicit clustered bounds does not have any areas that are extremely bad, regardless of depth complexity, and would thus perform more consistently. Most importantly, it’s worst case performance is much better than tiled.
Here we can see the impact of adding 2.5D culling to a tiled technique. While it helps in the discontinuity case (although does not reach clustered’s performance), it doesn’t help much or at all in a depth slope situation.
So the difference between tiled and clustered is that we pick a light list on a per-pixel basis instead of per-tile, depending on which cluster we fall within. Obviously though, in a lot of cases nearby pixels will choose the same light list, in particular neighbors within the same tile on a similar depth. If we visualize what light lists were chosen, we can see that there are a bunch of different paths taken beyond just the tile boundaries. A number of depth discontinuities from the foliage in front of the player gets clearly visible. This may seem like a big problem, but here we are only talking about fetching different data. This is not a problem for a GPU, it’s something they do all the time for regular texture fetches, and this is even much lower frequency than that.
The thing you might worry about though is divergent branches. However, despite fetching different light lists from pixel to pixel, the situation is not nearly as bad as you might expect from the previous picture. Chances are that the light lists look fairly similar. If you have one light lists with 5 lights and another with 5 lights (that are not necessarily the same as the other ones), branching will still be 100% coherent. You may pay a small overhead from the ideal when the lists have different light count, but that is typically going to be a relatively small overhead. In the worst-case scenario (no coherency at all), the amount of shading essentially boils down to what tiled shading has to shade.
The original paper [Olsson et. al 12] was written by academics, and naturally the direction of their research doesn’t match 100% with the requirements of the a game engine. We don’t have millions of tiny lights, but between hundreds and thousands of mostly artist placed lights, that are on a human scale. This meant that tight culling, so as to not add lights to more clusters than necessary, became more important to us. The higher-order clustering options the paper explored (and also largely rejected) were also something that we didn’t expect to work for us. Deriving the explicit cluster bounds was something that could be interesting, but we found that sticking to implicit bounds simplified the technique, while also allowing the light assignment to run on the CPU. This enables DX10 level GPU compatibility, which may be important given that (as of this writing) 24% on Steam are still on a DX10 GPU. In addition, this gives us scene independence. This means that we don’t need to know what the scene looks like to fill in the clusters, and this also allows us to evaluate light at any given point in space, even if it’s floating in thin air. This could be relevant for instance for ray-marching effects.

The paper only explored pointlights, whereas we need spotlights as well. We also needed a shadow solution, which the original paper also did not explore. However, Olsson et. al. has since continued their research and have now an interesting shadow approach made for clustered shading. We have however stuck with our own simpler approach. Finally, our games are massively large while still being played on human scale, resulting in a depth span from very near to very far, which required some extra fiddling to get rolling with clustered shading.
We are still using a deferred engine, but we could change to forward at any time should we decide that to be better. The important part is, however, that the transparency passes can now use the same lighting structure as the deferred passes, making it a unified lighting solution. Since we are still using deferred, and thus obviously have a complete depth buffer once we get to the deferred lighting pass, we could potentially use explicit bounds there. We still haven’t explored that opportunity, but it’s an option. It’s unclear if computing the explicit bounds, plus an extra round of culling, is going to be outweighed by potentially faster light evaluation.

Currently we are using 64x64 screen-space tiles, and 16 depth slices. This is most likely going to change, primarily because currently the tiles are currently fairly long and thin, and this is not optimal for a culling, in particular for spotlights. We have been experimenting with other setups, such as 128x128 and 32 depth slices. This created more cubical shaped clusters and helped with culling, which helped with culling, especially for spotlights. Another option we have considered, but not yet explored, is to not base it on pixel count, but simply divide the screen into a specific number of tiles regardless of resolution. This may reduce coherency on the GPU side somewhat in some cases, but would also decouple the CPU workload from the GPU workload and allow for some useful CPU side optimizations if the tile counts are known at compile time.
We are using exponential depth slicing, much like in the paper. There is nothing dictating that this is what we have to use, or for that matter that it is the best or most optimal depth slicing strategy; however, the advantage is that the shape of the clusters remain the same as we go deeper into the depth. On the other hand, clusters get larger in world space, which could potentially result in some distant clusters containing a much larger amount of lights. Depending on the game, it may be worth exploring other options.

Our biggest problem was that our depth ratio is massive, with near plane as close as 0.1m and far plane way out on the other side of the map, at 50,000m. This resulted in poor utilization of our limited depth slices, currently 16 of them. The step from one slice to the next is very large. Fortunately, in our game we don’t have any actual light sources beyond a distance of 500m. So we simply decided to keep our current distant light system for distances beyond 500m and limit the far range for clustering to that.

This improved the situation notably, but was still not ideal. We still burnt half of our slices on the first 7 meters from the camera. Given how our typical scenes look like, that’s likely going to be mostly empty space in most situations. So to improve the situation, we made the first slice special and made that go from near plane to an arbitrary visually tweaked distance, currently 5m. This gave us much better utilization.
This illustrates our distant light system, which has been around since Just Cause 2. In this screenshot there are likely no actual lights enabled since we’re far from civilization on top of a mountain, except perhaps our fake “night light” that slightly illuminates the area around the player at night to help game-play a bit in the darkness. Everything in the distance though, while representing actual artist placed lights, the actual light sources aren’t loaded at this distance. They are simply stored as a very compact list of point sprites, resident in memory at all time, and which is very cheap to render. We are at this point still using the same forward rendering solution here as in Just Cause 2, but one option now that we are using deferred is to actually compute real lighting under those sprites instead of just a putting a blob from a texture there.
Here’s the same system can be seen in effect in Just Cause 3. To the right of Rico’s arm and in the upper right corner we have instances of the distant light system, but also if you look down into the fog in a bottom left corner.
And in this screenshot you can also see the effect as a few lights in the distance in the upper center part of the screen.
This illustrates the benefit of the special near cluster. Less slices are wasted, and the cluster shapes aren’t quite as long and thin.
Given a screen position and a depth value (whether from a depth buffer or the rasterized depth in a forward pass) we start by looking up the cluster from a 3D texture. Each texel represents a cluster and its light list. The red channel gives us an offset to where the light list starts, whereas the green channel contains the light counts. The light lists are then stored in a tightly packed lists of indexes to the lights. The actual light source data is stored as arrays in a constant buffer.

All in all the data structure is very compact. In a typical artists lit scene it may be around 50-100kb of data to upload to the GPU every frame.
This shows the shader code for rendering with this data structure. The input is just the screen-space position and depth. This shows a deferred pass where depth comes from a texture, but in a forward pass the second line of code would simply use In.Position.z instead. Everything else would be identical, which shows how easily this technique adapts to either deferred or forward.

The ZParam.xy here contains the same parameters that you would use to compute a linear depth from a Z-buffer value, except I eliminated the division since that just becomes a negative under the logarithm, i.e. \( \log_2(1/(z*a+b)) = \log_2(z*(-a)+(-b)) \).
The light list could theoretically become huge. Say you have a total of 30*17*16 clusters at 1080p, and allow up to 256 lights per cluster, that would need 4MB, which with double-buffering (because it’s updated from the CPU) means you’ll need 8MB. Perhaps not a problem on next-gen, but hardly ideal, and who knows how many times these numbers will be bumped before you ship.

Normally, not every light affects every cluster in a scene. In fact, it’s extremely rare that you get even remotely close to that. So we constructed a somewhat plausible worst-case scenario with loads of large lights jammed in front of the player and recorded the max utilization ever encountered. Then multiplied up that for some extra margin. Even after that, the resulting buffer size we needed to allocate was far smaller. Naturally though, if you go down this path, it’s clearly important to add runtime assertions and warnings to make sure you don’t ever go above what you actually have allocated. Done correctly, at worst you would have artifacts for that extreme frame where a thousand nukes blew up in the player’s face.
Our light sources are typically artist placed, scaled for human environments in an outdoor world, so generally speaking from meters to tens of meters. So a light source generally intersects many clusters. The typical sphere-frustum tests that you can find online are not suitable for this sort of culling. They are made for view-frustum culling and based on the assumption that the frustum typically is much larger than the sphere, which is the opposite of what we have here. Typically they simply test sphere vs plane for each six planes of the frustum. This is conservative, but lets through spheres that aren’t completely behind any of the planes, such as in the frustum corners. The result you get is that green rectangle, or essentially a “cube” of clusters around the light. But that’s also the first thing we compute. We simply compute the screen-space and depth extents of the light analytically first, so this test doesn’t actually help anything at all after that.
Most frustum culling code is written with the scenario on the left in mind. We need to handle the scenario on the right.
One way to go about frustum culling is testing all planes, all edges and all vertices. This would work, but be too costly to outweigh the gains from fewer false positives. A fast, conservative but relatively tight solution is what we are looking for. There are many approaches that seem fitting, but there are also many complications, which has ultimately thrown many of our attempts into the garbage bin. One relatively straightforward approach is to cull against the cluster’s AABB. This is fast and gives fairly decent results, but it’s possible to do better.
Starting with the “cube” of clusters around the light, in our outer loop we iterate over the slices in z direction. We intersect the sphere with the slice where it is the widest. This results in a circle of a smaller radius than the original sphere, we thus continue in the y direction using a sphere of this smaller radius and the circle’s midpoint. In the center slice we simply proceed with the original sphere. We repeat this procedure in y and have an even smaller sphere. Then in the inner loop we do plane vs. sphere tests in x direction to get a strip of clusters to add the light to.

To optimize all the math we take advantage of the fact that in view-space, all planes will have components that are zero. A plane in the x direction will have zero y and offset, y direction has zero x and offset, and z-direction is basically only a z offset.

The resulting culling is somewhat tighter than a plain AABB test, and costs about the same. Where AABB culls around 15-25%, this technique culls around 20-30% from the “cube” of clusters.
Here’s the result visualized in 3D.
This shows the gist of the culling code.
For spotlights we begin by finding the “cube” of clusters around the light’s sphere, just like for pointlights, except this cube typically is much larger than necessary for a spotlight. However, this analytical test is cheap and goes a long way to limit the search space for following passes. Next we find a tighter “cube” simply by scanning in all six directions, narrowing it down by doing plane-cone tests. There is likely a neat analytical solution here, but this seemed non-trivial. Given that the plane scanning works fine and is cheap we haven’t really explored that path.

Note that our cones are sphere-capped rather than flat-capped. That’s because the light attenuation is based on distance (as it should), rather than depth. Sphere-capped cones also generally behave much better for wide angles and doesn’t become extremely large as flat-capped cones can get.
Finally, for the remaining “cube” of clusters we cull each cluster with a sphere-capped cone vs. bounding sphere test. For this to work well we have to have relatively cubical shaped clusters, otherwise the bounding sphere becomes way oversized. Overall this technique results in a moderately tight culling that is good enough for us so far, although there is room for some minor improvement.
Here’s the result visualized in 3D. Although our spotlights are sphere-capped, our debug visualization still draws them as flat-capped. That’s why it might look like it’s extending a bit outside the clusters.
Here’s the result with a handful of pointlights and spotlights enabled in a scene. The number of pointlights goes into red, and number of spotlights into green.
Classic deferred has the advantage that you can iterate light by light, and thus reuse resources such as shadow buffers in between. This saves some memory, which may be needed on current generation consoles. On PC and next-generation consoles this is not nearly as big a problem.

With the switch to clustered shading the cost of adding a new light to the scene is small. Artists can now be moderate “wasteful” without causing much problems performance-wise. This is not true for rasterizing shadow buffers. They remain expensive, and relatively speaking going to be more expensive going forward since it’s often a ROP-bound process, and ROPs aren’t getting scaled up nearly as much as ALU. So we still need to be a bit conservative about how many shadow casting lights we add to the scene.

An observation that was made is that artists often place very similar looking lights close to each other. In some cases it is to get a desired profile of a light, in which case the two lights may in fact be centered at the exact same point. But often it is motivated by the real world, such as two headlights on car. Some vehicles actually have ten or more lights, all pointing in the same general direction. Rendering ten shadow buffers for that may prove to be far too expensive.
Often it works just fine to share a single shadow buffer for these lights. While the shadow may be slightly off, this is usually not something that you will notice unless you are specifically looking for it. To make this work the shadow buffer is decoupled from lights and the light is assigned a shadow buffer and frustum from which to extract shadows. The shadow frustum has to be large enough to include all the different lights that uses it.
Given that we are doing the light assignment on the CPU, one may suspect that this will become a significant burden for the CPU. However, our implementation is fast enough to actually save us a bunch of CPU time over our previous solution. In a normal artist lit scene we recorded 0.1ms on one core for clustered shading. The old code supporting our previous forward pass for transparency that was still running in our system was still consuming 0.67ms for the same scene, a cost that we can now eliminate.

As of this writing, further optimizations have been made resulting in another 30-50% lower CPU cost than previously.
When we have nothing but the sun light in our scene, we incur a small overhead compared to classic deferred shading from looking up our empty light list and looping zero times. Once a light or two has been entered into the scene clustered shading is typically faster, and in regular artist lit scenes significantly so. Once we go to extreme artificial test cases with hundreds of lights sprinkled randomly in front of the player, clustered scales really well whereas classic deferred gets significantly slower. We have observed cases as large as 5x more expensive, whereas typically for heavy scenes it’s around 2x. The difference is generally about how slow we can make classic deferred, rather than how fast clustered can be, as the clustered performance stays quite consistent, whereas classic’s performance can very quite a lot depending on the scene.
We did some prototyping on a distance based clustering strategy instead of depth. While this allowed pointlights to be culled efficiently and exactly, this also made the cluster lookup slightly more expensive. The performance gain from an exact test was small enough that only with extreme workloads did we gain back what we lost from slower cluster lookup and we were hard-pressed to find a case where it ended up being faster in practice.

Another possible approach is clustering on view-space cascades. This would allow for exact AABB tests. One could argue that if you are going to test using an AABB, then you might just as well shape your clusters that way.

World-space clusters is another interesting option. While this would utilize the available clusters worse, the light distribution might match real world better. The other advantage is that you could evaluate light outside of the view frustum. This would allow for instance a reflection pass (such as a rear-view mirror) to use the same lighting structure for light evaluation.

There may be performance gains to be had if we consider the actual lights we have when clustering. For instance, we could tighten the cluster bounds if the most distant light active is closer than 500m, and the closest one more distant than 5m. This would allow for better cluster utilization.

Finally, a quick conservative reduction of the depth values in a shadow buffer could allow us to cull some clusters based on a conservative maximum-z value over some range. Whether this would result in any actual performance gains is unclear though.
Conclusions

- Clustered shading is practical for games
  - It's fast
  - It's flexible
  - It's simple
  - It opens up new opportunities
    - Evaluate light anywhere
    - Ray-trace your volumetric fog

What are you waiting for? Start writing your clustered shader today! 😊
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